

Spectrometer Calibration using Sparse Inverse Problems for Atmospheric Sounding

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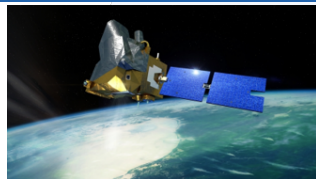
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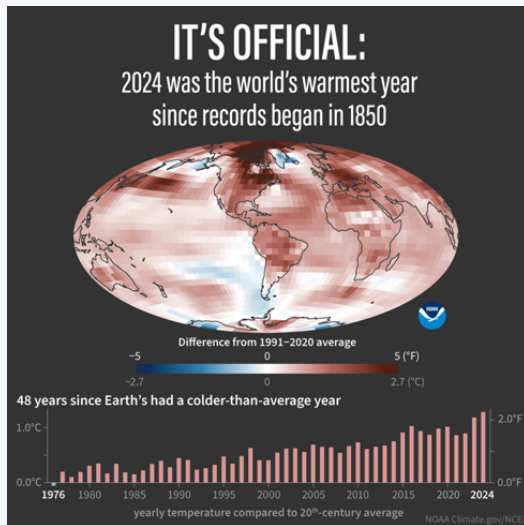
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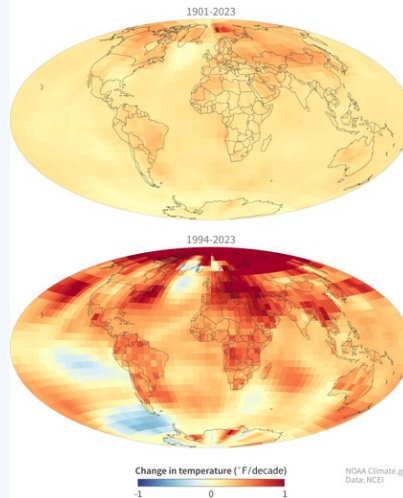


November 25, 2025

Climate change challenges

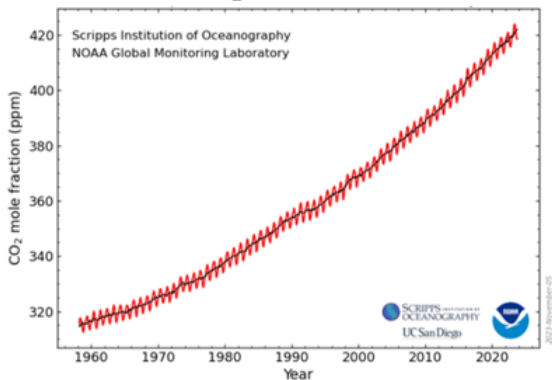


WARMING OVER PAST 30 YEARS IS MUCH FASTER THAN LONG-TERM TREND

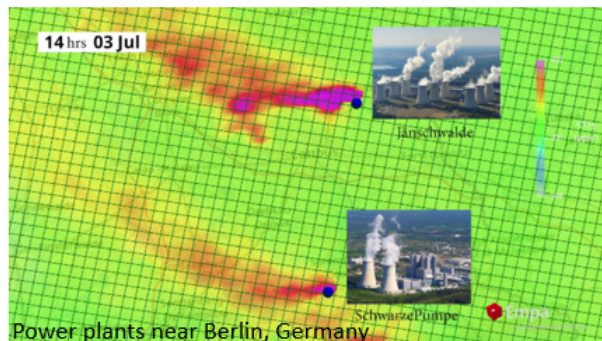


Climate change challenges

Atmospheric CO_2 at Mauna Loa Observatory



Simulated CO_2 at $2 \times 2 \text{ km}^2$ grid



Climate change challenges



Droughts and heat waves

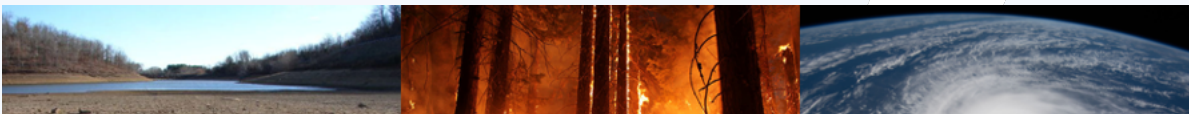


Intense wildfire seasons



Frequent natural disasters

Climate change challenges



Need to retrieve CO_2 concentrations

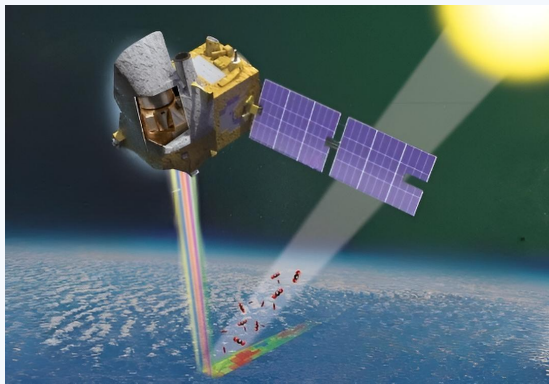


Droughts and heat waves

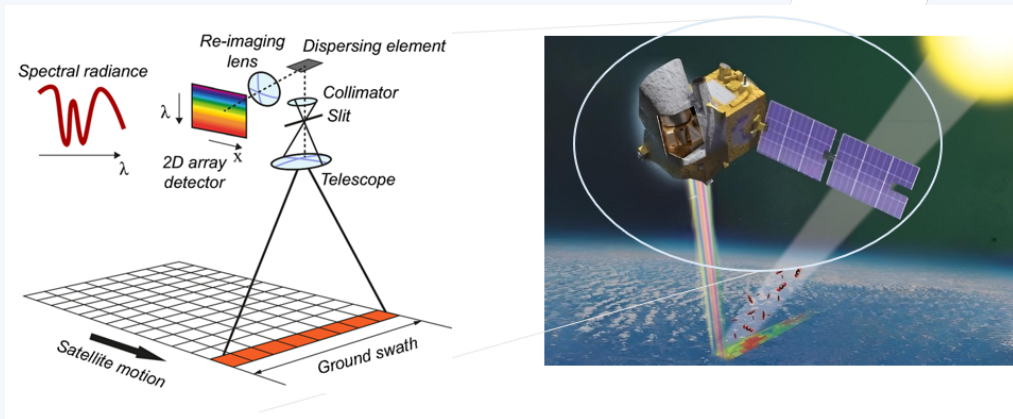
Intense wildfire seasons

Frequent natural disasters

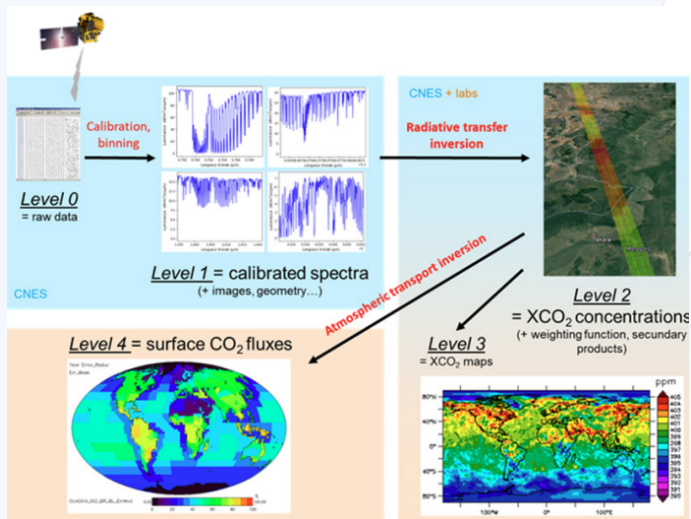
Monitoring Atmospheric CO_2 Concentrations via Remote Sensing



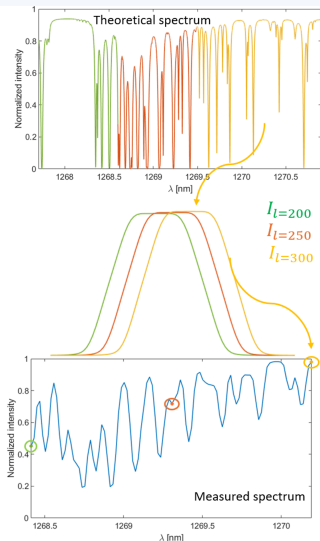
Monitoring Atmospheric CO_2 Concentrations via Remote Sensing



Spectrometer Calibration for Accurate CO_2 Concentration Measurements



Estimation of Instrument Spectral Response Functions (ISRFs)



An ill-posed problem

For each pixel $\ell = 1, \dots, N_\lambda$

- ▶ 1 wavelength λ_ℓ
- ▶ 1 observation $y_\ell \in \mathbb{R}^+$
- ▶ 1 ISRF $I_\ell \in \mathbb{R}^{N+1}$
- ▶ 1 nonlinear function f_ℓ
- ▶ 1 reference spectrum: $r_\ell = \pi(r, \lambda_\ell)$

$$y_\ell = f_\ell(\langle g(r_\ell), I_\ell \rangle) + \epsilon_\ell$$

known unknown unknown known unknown

Contributions of this Thesis

An ill-posed problem

For each pixel $\ell = 1, \dots, N_\lambda$

$$\underbrace{y_\ell}_{\text{known}} = \underbrace{f_\ell}_{\text{unknown}} \left(\underbrace{g}_{\text{unknown}} \left(\underbrace{r_\ell}_{\text{known}} \right), \underbrace{l_\ell}_{\text{unknown}} \right) + \epsilon_I$$

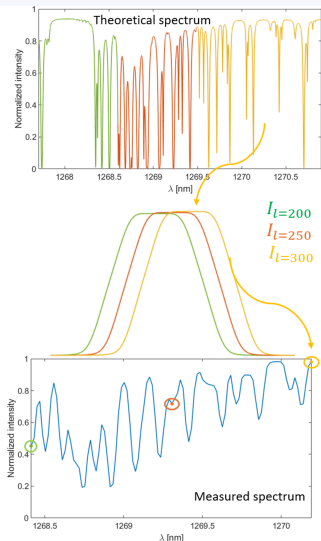
- ▶ **Linear inverse problem:** $y_\ell = \langle r_\ell, l_\ell \rangle + \epsilon_I$
 - ▶ New model: Sparse representation of ISRFs
- ▶ **Sparse nonlinear inverse problem:** $y_\ell = f_\ell(\langle r_\ell, l_\ell \rangle) + \epsilon_I$
 - ▶ Separable nonlinearities [Golub1973]
 - ▶ General model of nonlinearities
 - ▶ Dictionary preconditioning
- ▶ **Sparse nonlinear inverse problem with shifts:** $y_\ell = f_\ell(\langle g(r_\ell), l_\ell \rangle) + \epsilon_I$
 - ▶ Model of shifts

Outline

- 1 Context and Motivation
- 2 ISRF Estimation as an Inverse Problem
- 3 Sparse nonlinear inverse problem
- 4 Sparse nonlinear inverse problem with shifts
- 5 Conclusion & Perspectives

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ISRF Estimation as an Inverse Problem



An ill-posed problem

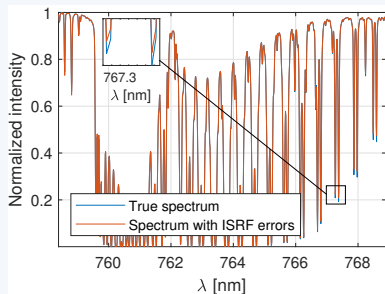
For each pixel $\ell = 1, \dots, N_\lambda$

- ▶ 1 wavelength λ_ℓ
- ▶ 1 observation $y_\ell \in \mathbb{R}^+$
- ▶ 1 ISRF $I_\ell \in \mathbb{R}^{N+1}$
- ▶ nonlinearities: functions f_ℓ and g
- ▶ 1 reference spectrum: $r_\ell = \pi(r, \lambda_\ell)$

$$y_\ell = f_\ell \left(g \left(r_\ell \right), I_\ell \right) + \epsilon_\ell$$

known unknown unknown known unknown

ISRF Estimation as an Inverse Problem



An ill-posed problem

For each pixel $\ell = 1, \dots, N_\lambda$

- ▶ 1 wavelength λ_ℓ
- ▶ 1 observation $s_\ell \in \mathbb{R}^+$
- ▶ 1 ISRF $l_\ell \in \mathbb{R}^{N+1}$
- ▶ 1 reference spectrum: $r_\ell = \pi(r, \lambda_\ell)$

$$\underbrace{s_\ell}_{\text{known}} = \langle \underbrace{r_\ell}_{\text{known}}, \underbrace{l_\ell}_{\text{unknown}} \rangle + \epsilon_l$$

ISRF Estimation as an Inverse Problem

Assumption: ISRFs $\mathbf{l}_\ell$ vary little from ℓ to $\ell + 1$

Sliding window $\mathcal{W}(\lambda_\ell)$ of $L + 1$ measurements around λ_ℓ

▶ $\mathbf{R}_\ell = \Pi(\mathbf{r}, \mathcal{W}(\lambda_\ell)) = [\mathbf{r}_{\ell-\frac{1}{2}}, \dots, \mathbf{r}_{\ell+\frac{1}{2}}] \in \mathbb{R}^{(L+1) \times (N+1)}$

▶ $\mathbf{s}_\ell = [s_{\ell-\frac{1}{2}}, \dots, s_{\ell+\frac{1}{2}}] \in \mathbb{R}^{L+1}$

Observation model (matrix form)

$$\mathbf{s}_\ell = \mathbf{R}_\ell \mathbf{l}_\ell$$

Assumptions about ISRF $\mathbf{l}_\ell$: positivity

Parametric models

Gaussian model ¹

$$I_{\ell, \beta_G}(x) = A_G \exp \left(-\frac{(\lambda_\ell - x - \mu_G)^2}{2\sigma_G^2} \right), \quad x \in \Delta$$

Δ wavelength grid, $\beta_G = (A_G, \mu_G, \sigma_G^2)^T$ - parameters

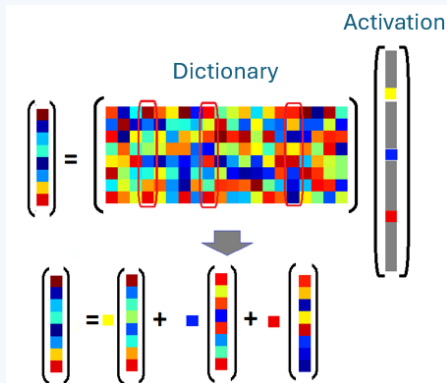
Super-Gaussian model ¹

$$I_{\ell, \beta_{SG}}(x) = A_{SG} \exp \left(-\left| \frac{\lambda_\ell - x - \mu_{SG}}{w_{SG}} \right|^{k_{SG}} \right), \quad x \in \Delta$$

$\beta_{SG} = (A_{SG}, \mu_{SG}, w_{SG}, k_{SG})^T$ - parameters

¹Beirle S. et al. (2017) – Parameterizing the instrumental spectral response function and its changes by a super-Gaussian and its derivatives, *Atmos. Meas. Tech.*

Proposed non-parametric model



Sparse representation of ISRFs (SPIRIT)

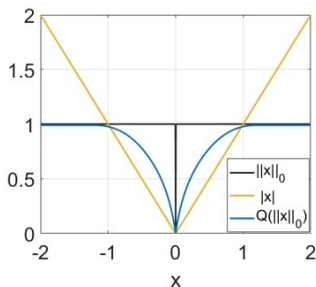
- ▶ Dictionary $\Phi \in \mathbb{R}^{(N+1) \times N_D}$
(SVD of representative ISRFs)

$$I_\ell \approx I_\ell^K = \Phi \alpha_\ell = \sum_{k=1}^K \alpha_k \Phi_{\gamma_k}$$

$$s_\ell \approx R_\ell I_\ell^K = \Psi_\ell \alpha_\ell$$

where $\Psi_\ell = R_\ell \Phi \in \mathbb{R}^{(L+1) \times N_D}$

ISRF Estimation as a sparse inverse problem



Optimization problem with ℓ_0 norm ²

$$\arg \min_{\alpha_\ell} \|\mathbf{s}_\ell - \Psi_\ell \alpha_\ell\|_2^2$$
$$s.t. \|\alpha_\ell\|_0 \leq K$$

Relaxation via ℓ_1 regularization (LASSO)

$$\arg \min_{\alpha_\ell} \|\mathbf{s}_\ell - \Psi_\ell \alpha_\ell\|_2^2 + \mu \|\alpha_\ell\|_1$$

Relaxation via quadratic envelope regularization ³

$$\arg \min_{\alpha_\ell} \|\mathbf{s}_\ell - \Psi_\ell \alpha_\ell\|_2^2 + Q_\gamma(\mu \|\alpha_\ell\|_0)(\alpha_\ell), \gamma > 0$$

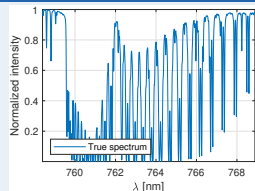
²Mallat, S., Zhang, Z. (1993) - Matching pursuits with time-frequency dictionaries. *IEEE Trans. Signal Process.*

³Carlsson M. (2019) – On convex envelopes and regularization of non-convex functionals, *J. Optim. Theory Appl.*

Experimental Setup and Evaluation Metrics

Data description

- ▶ Reference spectra: 4A/OP radiative transfer software
- ▶ Simulated ISRFs for the MicroCarb spectrometer
- ▶ Measured spectra: $s(\lambda_l) = (r * I_l)(\lambda_l) = \int_{-\infty}^{\infty} r(\lambda_l - u) I_l(u) du$

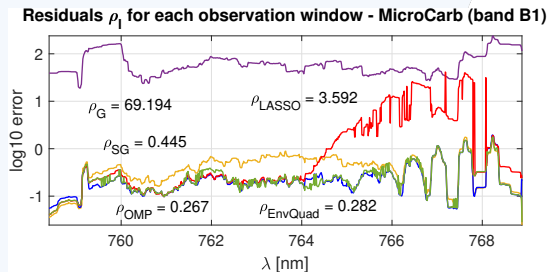


Metrics

- ▶ ISRF approximation errors: $E_\ell = \sum_{n=-N/2}^{N/2} |I_\ell(n\Delta) - \hat{I}_\ell(n\Delta)|$
- ▶ Residual errors: $\rho = \frac{1}{N_\lambda} \sum_{\ell=1}^{N_\lambda} \|s_\ell - r_\ell \hat{I}_\ell\|_2^2$

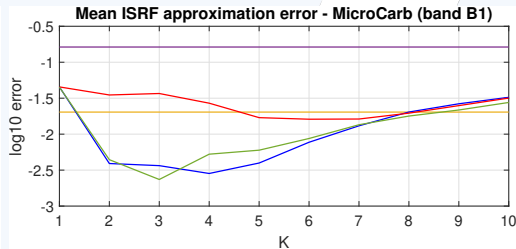
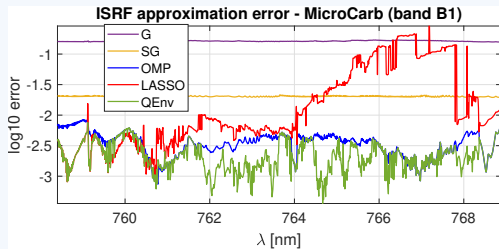
Numerical results

— G
— SG
— OMP
— LASSO
— QEnv



$$K = 3$$

Numerical results



ISRF estimation error ($K = 3$) and mean ISRF estimation error

- ▶ Poor performance with parametric models (0.1s)
- ▶ Limited performance of the LASSO algorithm (0.01s)
- ▶ Best performance achieved with quadratic envelopes (1s)
- ▶ OMP: best trade-off between computation time and estimation performance (0.001s)

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Sparse nonlinear inverse problem

Forward model

$$\underbrace{y_\ell}_{\text{known}} = \underbrace{f_\ell}_{\text{unknown}} \left(\underbrace{g}_{\text{unknown}} \left(\underbrace{r_\ell}_{\text{known}} \right), \underbrace{l_\ell}_{\text{unknown}} \right) + \epsilon_l \quad \ell = 1, \dots, N_\lambda$$

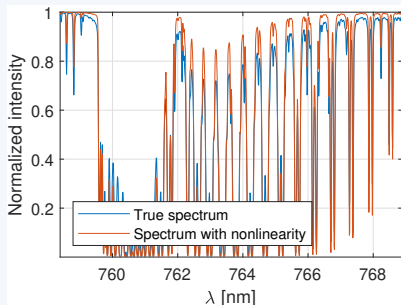
$$l_\ell = \Phi \alpha_\ell, \quad \alpha_\ell \text{ sparse}$$

Sparse nonlinear inverse problem

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$$l_\ell = \Phi \alpha_\ell, \quad \alpha_\ell \text{ sparse}$$



- ▶ Parameterized nonlinearities: function f_θ
 - ▶ Spectrometer calibration
 - ▶ Hyperspectral imaging
 - ▶ Epidemiology

Description of nonlinearities

Parameterized nonlinearities

- ▶ f_θ bijective
- ▶ Separable nonlinearity⁴: $f_\theta(z) = \sum_{p=0}^P g_p(z)\theta_p \quad \forall z \in \mathbb{R}^+$
 - ▶ Polynomials⁵
- ▶ General model of nonlinearities^{6,7}
 - ▶ $f_\theta = \theta_1 x^{\theta_2}$
 - ▶ $f_\theta = \tanh(\theta_1 x) + \theta_2 x$

⁴Golub, G. H., Pereyra V. (2003) - Separable nonlinear least squares: the variable projection method and its applications, *Inverse Probl.*

⁵Altmann Y. et al. (2012), Supervised nonlinear spectral unmixing using a postnonlinear mixing model for hyperspectral imagery, *IEEE Trans. Image Process.*

⁶Taghvaei A. et al. (2020), Fractional SIR epidemiological models, *Sci. Rep.*

⁷Babaie-Zadeh M. et al. (2001), Blind separating convolutive post non-linear mixtures, *Proc. 3rd ICA Workshop.*

Iterative optimization method

Extended Invariance Principle (EXIP) with sparsity

- ▶ Problem formulation with sparsity: $\arg \min_{\theta, \alpha_\ell} \|\mathbf{y}_\ell - f_\theta \circ (\Psi_\ell \alpha_\ell)\|_2^2 + \mu \|\alpha_\ell\|_0$
- ▶ Resolution method: An iterative approach
 - ▶ Step 1: Estimation of θ : $\arg \min_{\theta} \|\mathbf{y}_\ell - f_\theta \circ (\Psi_\ell \alpha_\ell)\|_2^2$
 - ▶ Step 2: EXtended Invariance Principle (EXIP)⁸ for estimating α_ℓ :
 $\beta_\ell = g(\alpha_\ell) = \Psi_\ell \alpha_\ell$

$$\arg \min_{\alpha_\ell} \left[\hat{\beta}_\ell - g(\alpha_\ell) \right]^T R''_{\beta_\ell}(\hat{\beta}_\ell) \left[\hat{\beta}_\ell - g(\alpha_\ell) \right] + \mu \|\alpha_\ell\|_0$$

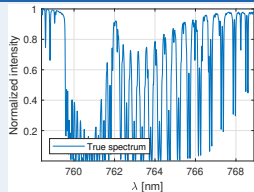
$$\text{where } R_{\beta_\ell}(\beta_\ell) = \|\mathbf{y}_\ell - f_\theta \circ \beta_\ell\|_2^2 \text{ and } R''_{\beta_\ell}(\hat{\beta}_\ell) = \nabla^2 R_{\beta_\ell}(\hat{\beta}_\ell)$$

⁸Stoica P. & Soderstr T. (1989) - On reparametrization of loss functions used in estimation and the invariance principle, *Signal Process.*

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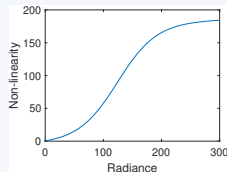


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- ▶ Residual errors: $\rho = \frac{1}{N_\lambda} \sum_{\ell=1}^{N_\lambda} \|s_\ell - r_\ell \hat{I}_\ell\|_2^2$
- ▶ Estimation errors for the parameters of the nonlinearities:
$$r_{NL} = \frac{1}{Q} \sum_{i=1}^Q (f_\theta(x_i) - f_{\hat{\theta}}(x_i))^2, x_i \in \mathcal{D}$$

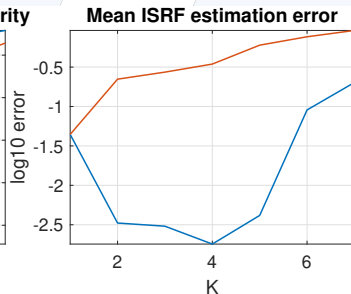
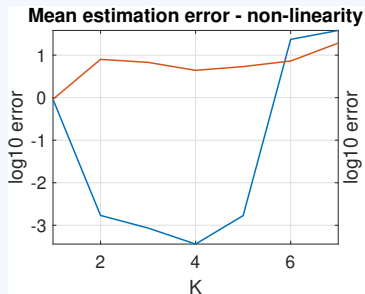
Numerical results - ISRF estimation

- Nonlinearity: $f_{\theta}(z) = \frac{\theta_1}{1+\exp(-\theta_2 z)} + \theta_2 z$



- Measured spectrum divided into 14 channels of $Q = 70$ wavelengths SNR=55dB
- 1 ISRF to estimate per wavelength

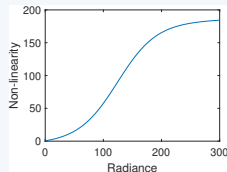
— EXIP-OMP
— OMP



Competitive performance with EXIP / OMP

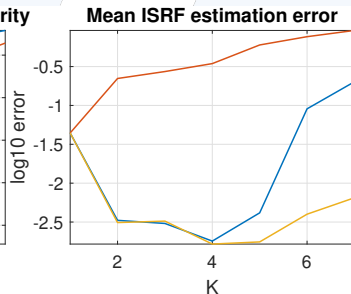
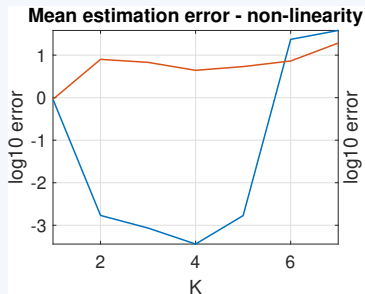
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— EXIP-OMP
— OMP
— OMP-Linear case



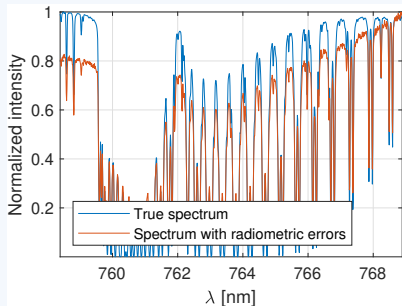
Competitive performance with EXIP / OMP

Radiometric calibration using multiple reference spectra

MicroCarb spectrometer

$$\underbrace{y_\ell}_{\text{known}} = \underbrace{f_{\theta_\ell}}_{\text{unknown}} \left(\langle \underbrace{r_\ell}_{\text{known}}, \underbrace{l_\ell}_{\text{unknown}} \rangle \right) + \epsilon_l \quad \ell = 1, \dots, N_\lambda$$

θ_ℓ - parameter vector of pixel ℓ



Radiometric errors: (functions f_{θ_ℓ})

- ▶ Linear gains
 - ▶ Dark current
 - ▶ Nonlinear gains
-
- ▶ 1 function f_{θ_ℓ} per pixel ℓ
 - ▶ Q reference spectra $r^{(q)}$

Radiometric Calibration Using Multiple Reference Spectra

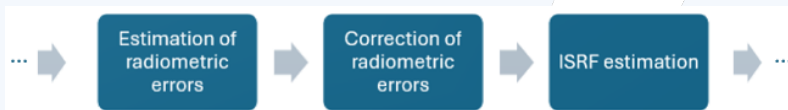
Joint estimation of ISRF and radiometric errors (SPIRITUAL)

- ▶ Polynomial model:

$$y_{\ell,q} = f_{\theta_{\ell}}(s_{\ell,q}) = \sum_{p=0}^P \theta_p^{\ell} (s_{\ell,q})^p = \sum_{p=0}^P \theta_p^{\ell} \left((\mathbf{r}_{\ell}^{(q)})^T \mathbf{I}_{\ell} \right)^p, \quad q = 1, \dots, Q$$

- ▶ Iterative estimation:

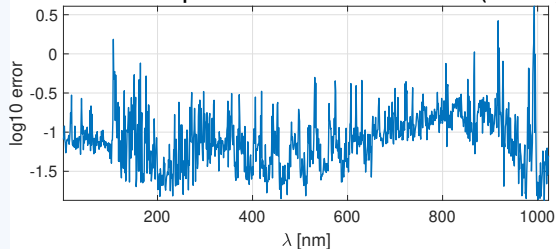
- ▶ **Step 1:** Estimate nonlinearity parameters θ_{ℓ} using Q reference spectra
- ▶ **Step 2:** Estimate ISRFs via sliding window with all reference spectra



Numerical results - Radiometric calibration using multiple reference spectra

— SPIRITUAL

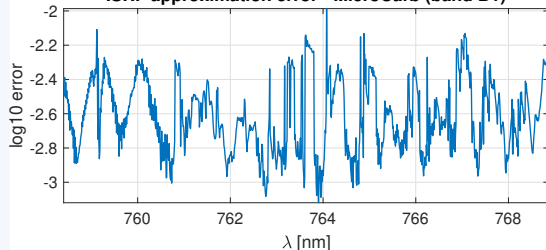
Radiometric response estimation error - MicroCarb (band B1)



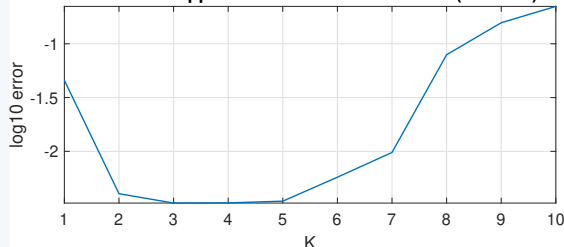
$K=4, P=3, Q=11$

Competitive performance with SPIRITUAL

ISRF approximation error - MicroCarb (band B1)

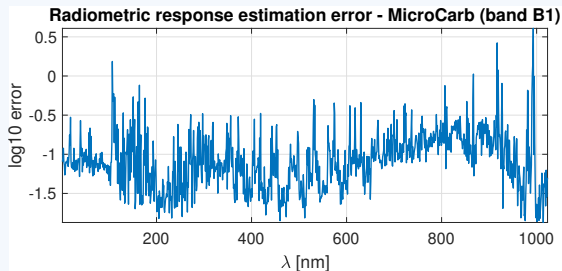


Mean ISRF approximation error - MicroCarb (band B1)



Numerical results - Radiometric calibration using multiple reference spectra

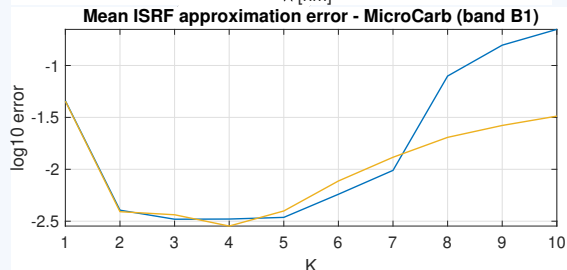
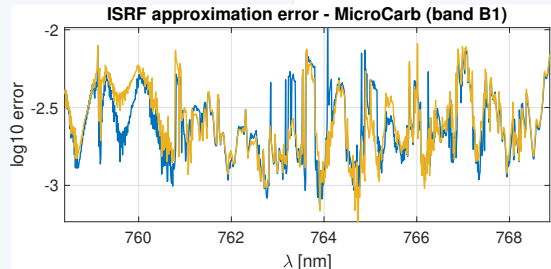
— SPIRITUAL
— OMP-Linear Case



$K=4, P=3, Q=11$

Competitive performance with SPIRITUAL

TēSA



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Sparse nonlinear inverse problem with shifts

Forward model

$$\begin{array}{ccccccc}
 \underbrace{y_\ell}_{\text{known}} & = & \underbrace{f_\ell}_{\text{unknown}} & \left(\underbrace{g}_{\text{unknown}} \left(\underbrace{r_\ell}_{\text{known}} \right), \underbrace{I_\ell}_{\text{unknown}} \right) + \epsilon_\ell & \ell = 1, \dots, N_\lambda \\
 \\
 I_\ell = \Phi \alpha_\ell, & \alpha_\ell \text{ sparse}
 \end{array}$$

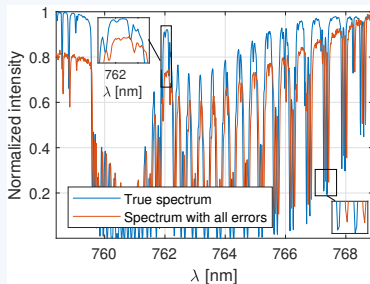
Sparse nonlinear inverse problem with shifts

Forward model

$$y_\ell = f_\ell (g(r_\ell), I_\ell) + \epsilon_\ell \quad \ell = 1, \dots, N_\lambda$$

known unknown unknown known unknown

$$I_\ell = \Phi \alpha_\ell, \quad \alpha_\ell \text{ sparse}$$



Parameterized nonlinearities:

- ▶ f_θ
 - ▶ Spectrometer calibration
 - ▶ Hyperspectral imaging
- ▶ shifts g_c
 - ▶ Incorrect values of λ_ℓ
 - ▶ Doppler effect

Sparse nonlinear inverse problem with shifts

Joint estimation of sparse vectors, nonlinearity and shifts

► Shift: $\lambda'_\ell = \lambda_\ell + g_{\mathbf{c}}(\ell)$

► Joint estimation problem: $\arg \min_{\boldsymbol{\alpha}, \boldsymbol{\theta}, \mathbf{c}} \sum_{\ell=1}^{N_\lambda} \|\mathbf{y}_\ell - f_{\boldsymbol{\theta}} \circ (\mathbf{R}_\ell(\mathbf{c}) \Phi \boldsymbol{\alpha}_\ell)\|_2^2 + \mu \sum_{\ell=1}^{N_\lambda} \|\boldsymbol{\alpha}_\ell\|_0$
 $\mathbf{R}_\ell(\mathbf{c}) \triangleq \Pi(\mathbf{r}, \mathcal{W}(\lambda_\ell + g_{\mathbf{c}}(\ell)))$

Sparse nonlinear inverse problem for MicroCarb

Problem in presence of radiometric errors and spectral shifts

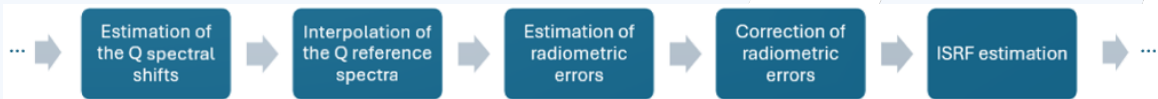
- ▶ Q reference spectra $\mathbf{r}^{(q)}$
- ▶ 1 function f_{θ_ℓ} per pixel ℓ
- ▶ 1 function g_{c_q} per reference spectrum q
- ▶ Forward model assuming polynomial nonlinearities for pixel ℓ , reference spectrum q :

$$y_{\ell,q} = \sum_{p=0}^P \theta_p^\ell \left((\mathbf{r}_\ell^{(q)}(\mathbf{c}_q))^T \mathbf{l}_\ell \right)^p,$$
$$\mathbf{r}_\ell^{(q)}(\mathbf{c}_q) \triangleq \pi \left(\mathbf{r}^{(q)}, \lambda_\ell + \sum_{m=0}^M c_m^q \left(\frac{\ell}{\ell_{\max}} \right)^m \right)$$

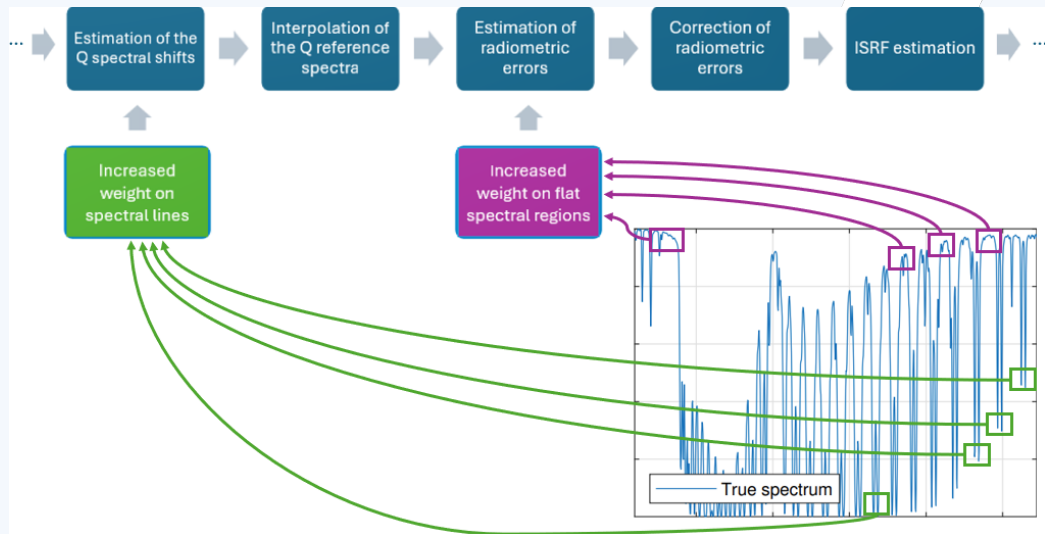
Sparse nonlinear inverse problem for MicroCarb

Iterative estimation in presence of radiometric errors and spectral shifts - ASPIRIT

- ▶ Estimate spectral shift parameters c_q for the Q reference spectra
- ▶ Estimate radiometric parameters θ_ℓ for the N_λ pixels
- ▶ Estimate ISRFs via sliding window with all reference spectra

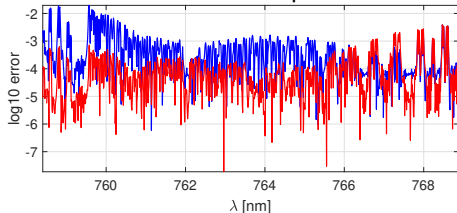


Radiometric and spectral calibration - Weighted ASPIRIT algorithm

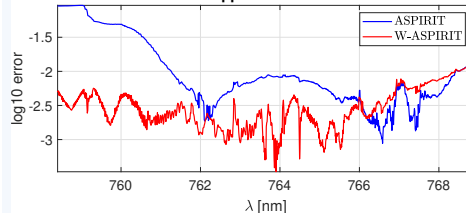


Numerical results - In presence of radiometric errors and spectral shifts

Differences between normalized measured spectrum and its reconstructions



ISRF approximation error



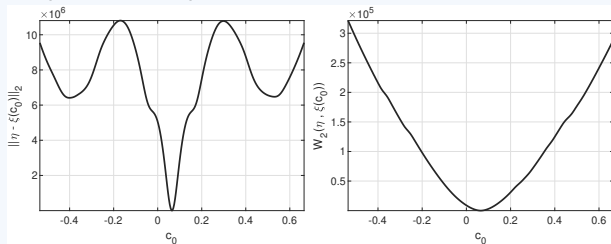
Measurement reconstruction error and ISRF estimation error versus λ for
 $K = 4$, $Q = 11$, $P = 3$ and $M = 3$

Promising results with the proposed method
Best performance obtained when using weights

Optimal Transport with Wasserstein distance

- ▶ Joint estimation problem: $\arg \min_{\alpha, \theta_\ell, c} \sum_{\ell=1}^{N_\lambda} \|y_\ell - f_{\theta_\ell} \circ (R_\ell(c)\Phi\alpha_\ell)\|_2^2 + \mu \sum_{\ell=1}^{N_\lambda} \|\alpha_\ell\|_0$

- ▶ Optimal Transport



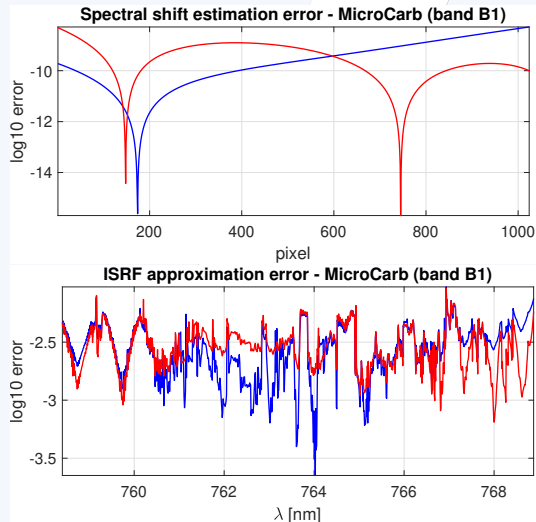
$$\arg \min_c \sum_{\ell=1}^{N_\lambda} W_2(y_\ell, R_\ell(c)I_\ell)$$

Distances between measurements s_ℓ and model $R_\ell(c)\Phi\alpha_\ell$ versus shift value

Numerical results - In presence of spectral shifts

— Norme ℓ_2 - $\delta_{\max} = 3$
— OT - $\delta_{\max} = 3$

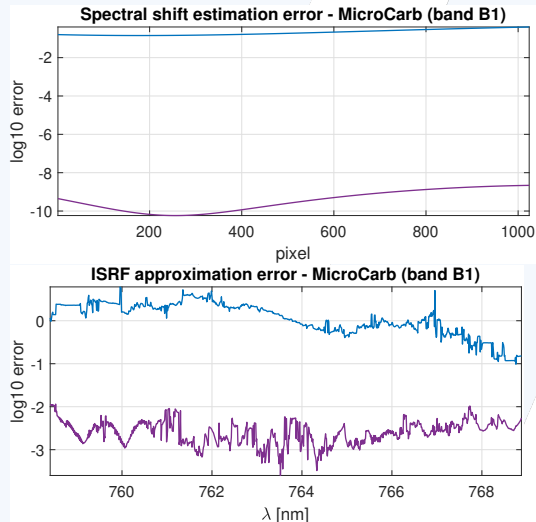
Spectral shift estimation error and
ISRF estimation error versus λ for
 $K = 4$, $M = 3$ and SNR= 55dB



Numerical results - In presence of spectral shifts

— Norme ℓ_2 - $\delta_{\max} = 38$
— OT - $\delta_{\max} = 38$

Spectral shift estimation error and
ISRF estimation error versus λ for
 $K = 4$, $M = 3$ and SNR= 55dB



- 1 Context and Motivation
- 2 ISRF Estimation as an Inverse Problem
- 3 Sparse nonlinear inverse problem
- 4 Sparse nonlinear inverse problem with shifts
- 5 Conclusion & Perspectives**

Conclusion

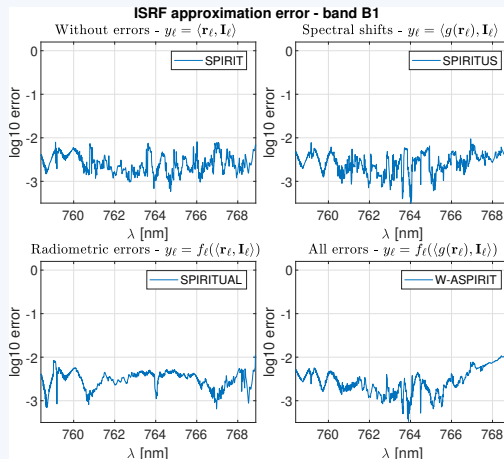
- ▶ Sparse inverse problem for spectrometer calibration
 - ▶ Best performance with OMP
- ▶ Nonlinear sparse inverse problem models with different nonlinearities
 - ▶ Competitive performance using EXIP and OMP
- ▶ Impact of Shifts
 - ▶ Importance of including weights

Conclusion

- ▶ Enhanced calibration strategies for atmospheric spectrometers (MicroCarb, CO2M)
- ▶ Sparse inverse problem models enabling high-accuracy recovery of ISRFs, spectral shifts, and radiometric nonlinearities
- ▶ Framework applicable beyond the targeted missions

Conclusion & Perspectives

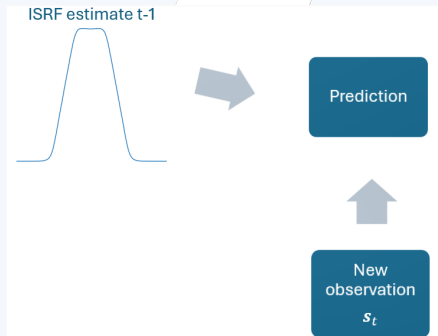
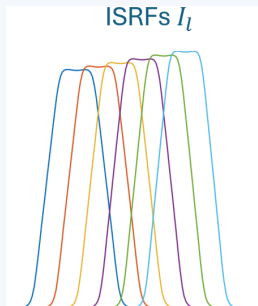
- Moving toward operational integration for space instruments



Perspectives

Alternative ISRF Estimation Models

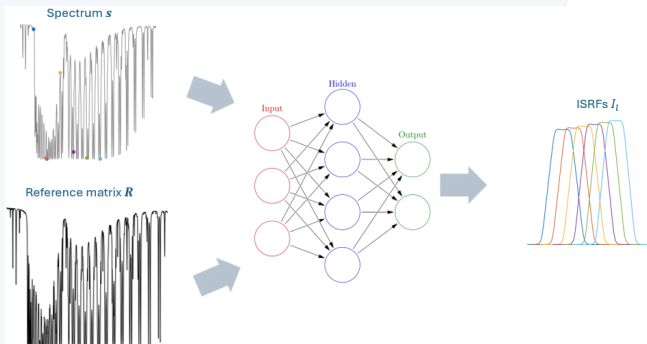
- ▶ State-space models
- ▶ Possible use of Kalman filter



Perspectives

Potential use of Machine Learning

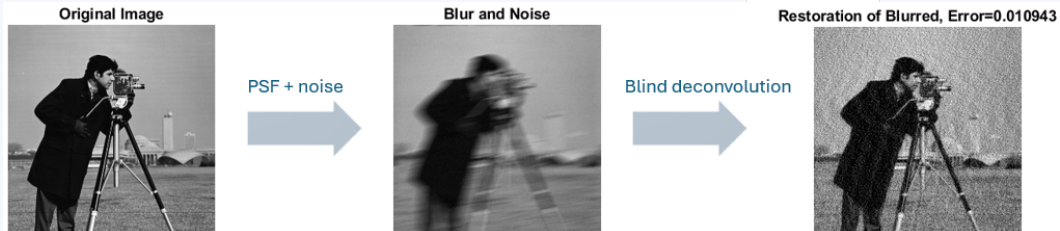
- ▶ Neural Network to estimate ISRF
- ▶ Learn nonlinear behaviors from data



Perspectives

Imperfect Knowledge of Reference Spectra

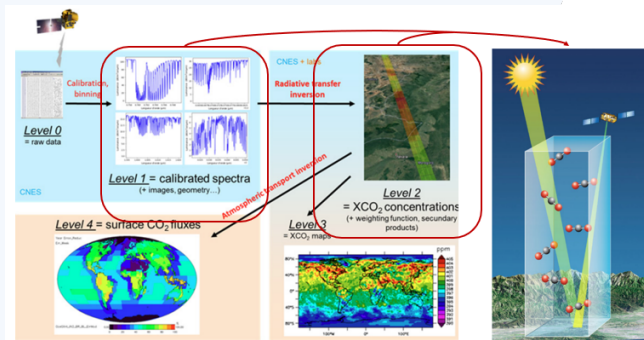
- ▶ ISRF estimation with reference spectrum uncertainties
- ▶ Blind deconvolution



Perspectives

Toward integration within CO₂ Retrieval Algorithms

- ▶ Use the proposed inversion methods to retrieve CO₂
- ▶ Perform joint calibration (level 1) and gas retrieval (level 2)



Publications

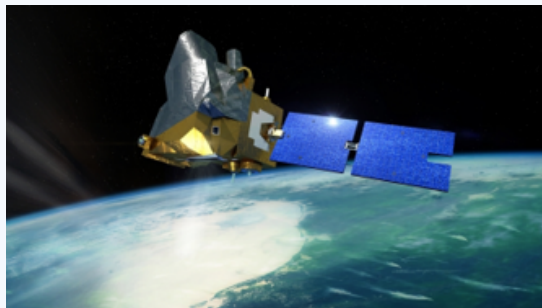
Journal papers

1. Jihanne El Haouari, Jean-Michel Gaucel, Christelle Pittet, Jean-Yves Tournet and Herwig Wendt, In-Flight Estimation of Instrument Spectral Response Functions Using Sparse Representations, Atmospheric Measurement Techniques, 2024
2. Jihanne El Haouari, Jean-Michel Gaucel, Christelle Pittet, Jean-Yves Tournet and Herwig Wendt, Solving parametric nonlinear inverse problems through sparse modeling, IEEE Transactions on Signal Processing, 2025, in preparation

Conference papers

1. Jihanne El Haouari, Jean-Yves Tournet, Herwig Wendt, Christelle Pittet, Jean-Michel Gaucel, Approximation et Estimation de Réponses Spectrales d'Instruments, GRETSI, Grenoble, France, August 28- September 1, 2023
2. Jihanne El Haouari, Jean-Michel Gaucel, Christelle Pittet, Jean-Yves Tournet and Herwig Wendt, Estimation of Instrument Spectral Response Functions using Sparse Representation in a Dictionary, Proc. IEEE IGARSS, Athens, Greece, July 7-12, 2024
3. Jihanne El Haouari, Jean-Michel Gaucel, Christelle Pittet, Jean-Yves Tournet and Herwig Wendt, Estimation of Instrument Spectral Response Functions in Presence of Radiometric Errors, Proc. EUSIPCO, Lyon, France, Aug. 26-30, 2024.
4. Jihanne El Haouari, Jean-Michel Gaucel, Christelle Pittet, Jean-Yves Tournet and Herwig Wendt, On-ground and in-flight estimation of instrument spectral responses in the presence of measurement errors, ICSO2024, Antibes Juan-Les-Pins, Oct. 21-25, 2024
5. Jihanne El Haouari, Marcus Carlsson, Jean-Yves Tournet, Herwig Wendt, Jean-Michel Gaucel and Christelle Pittet, Estimating Instrument Spectral Response Functions Using Sparse Representations and Quadratic Envelopes, ICASSP, Hyderabad, India, Apr. 06-11, 2025
6. Jihanne El Haouari, Filip Elvander, Jean-Michel Gaucel, Christelle Pittet, Jean-Yves Tournet and Herwig Wendt, Joint ISRF and Spectral Shift Estimation for Spectrometer Calibration using Optimal Transport, Proc. EUSIPCO, Palermo, Italy, Sept. 8-12, 2025

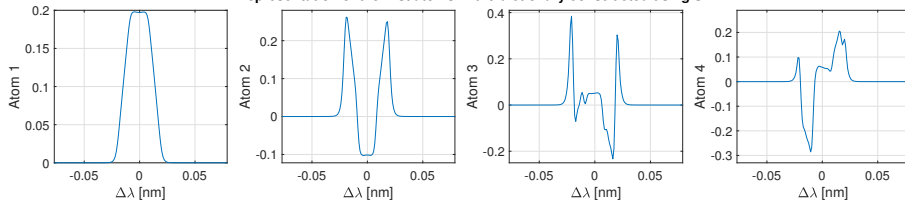
Thank you for your attention



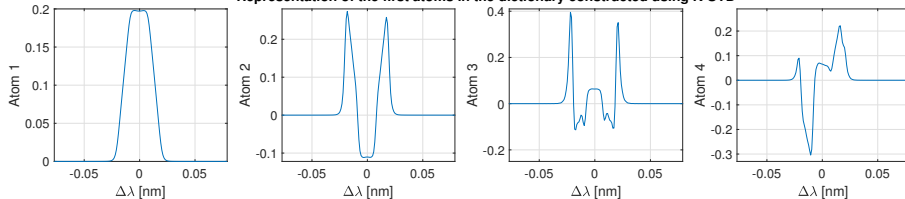
I look forward to your questions

Atoms of the dictionaries

Representation of the first atoms in the dictionary constructed using SVD



Representation of the first atoms in the dictionary constructed using K-SVD



Robustness to noise

Mean ISRF approximation error and mean residual as a function of SNR and method (G, SG, OMP, LASSO, SVD, K-SVD) for MicroCarb (band B1) —
in blue: ISRF estimation errors under the mission criterion; in bold: best results

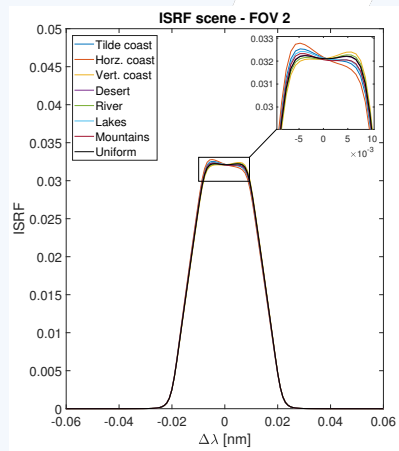
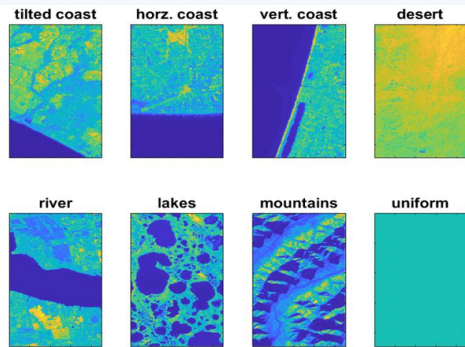
		ISRF approximation error (%)						Residual					
Instrument	/ SNR	G	SG	OMP SVD	OMP K-SVD	LASSO SVD	LASSO K-SVD	G	SG	OMP SVD	OMP K-SVD	LASSO SVD	LASSO K-SVD
MicroCarb bande B1	20 dB	16.28	3.39	4.58	4.37	14.38	14.23	185.5	116.4	112.3	112.4	112.7	112.9
	40 dB	16.27	2.04	0.54	0.56	2.05	2.37	70.2	1.56	1.32	1.32	1.53	1.70
	55 dB	16.27	2.03	0.29	0.33	1.33	1.66	69.21	0.43	0.23	0.24	0.38	0.61
	80 dB	16.27	2.03	0.28	0.32	1.27	1.68	69.2	0.39	0.20	0.20	0.34	0.60
	120 dB	8.03	7.79	0.51	0.55	0.83	0.73	45.07	37.32	1.11	1.13	1.14	1.13

Robustness to noise in presence of radiometric errors

Mean ISRF approximation error for MicroCarb (band B1) — in blue: ISRF estimation errors under the mission criterion

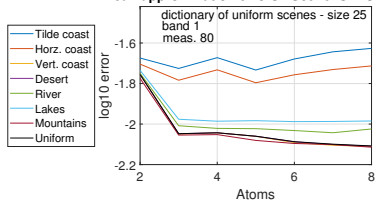
		ISRF approximation error (%)
Instrument / SNR		SPIRITUAL
MicroCarb bande B1	20 dB	1.01
	40 dB	0.29
	55 dB	0.27
	80 dB	0.27

ISRF sensitivity to the scene

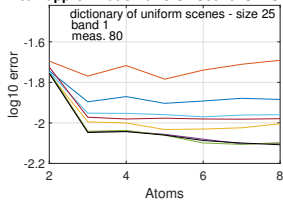


ISRF sensitivity to the scene

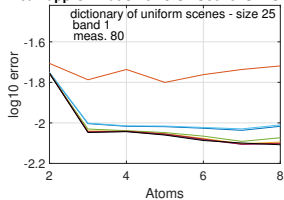
Mean approximation errors - Scene ISRFs FOV 1



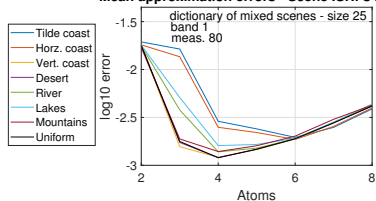
Mean approximation errors - Scene ISRFs FOV 2



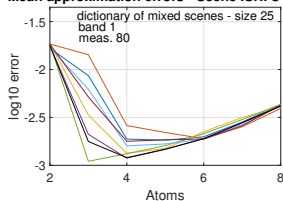
Mean approximation errors - Scene ISRFs FOV



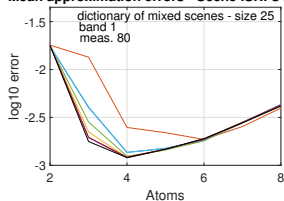
Mean approximation errors - Scene ISRFs FOV 1



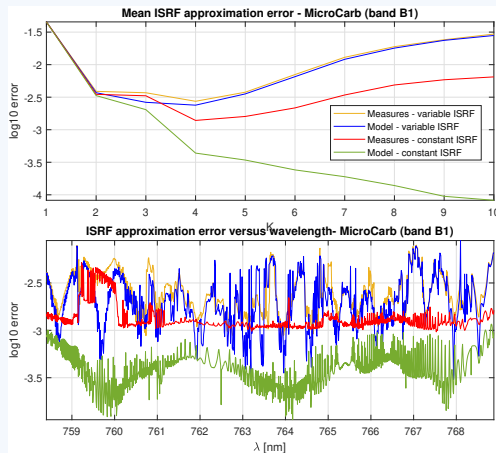
Mean approximation errors - Scene ISRFs FOV 2



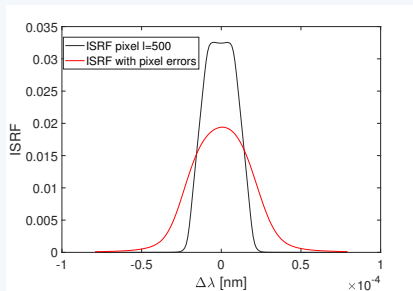
Mean approximation errors - Scene ISRFs FOV



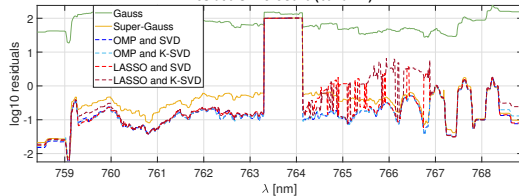
Interpolation-induced Mismatch



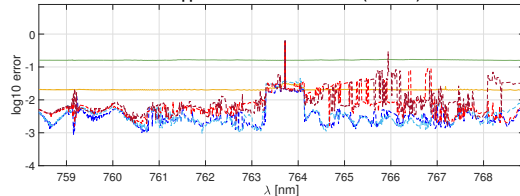
Uncertainties about one ISRF



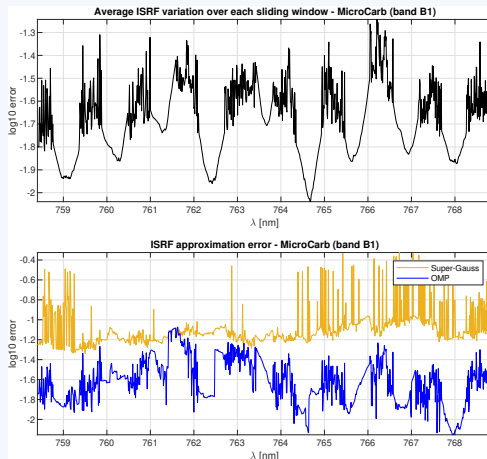
Residuals MicroCarb (band B1)



ISRF approximation error MicroCarb (band B1)

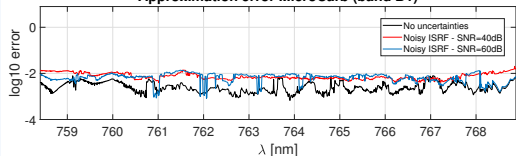


Uncertainties about the sliding window

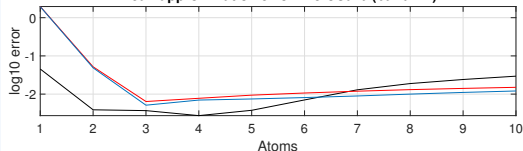


Uncertainties about the dictionary and the reference spectrum

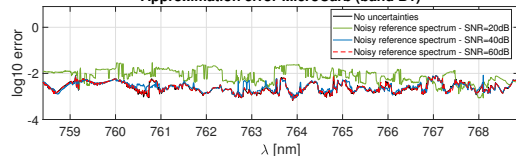
Approximation error MicroCarb (band B1)



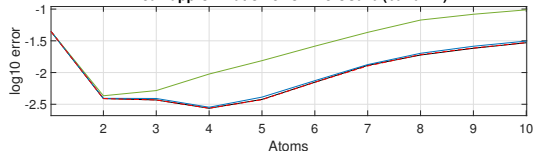
Mean approximation error MicroCarb (band B1)



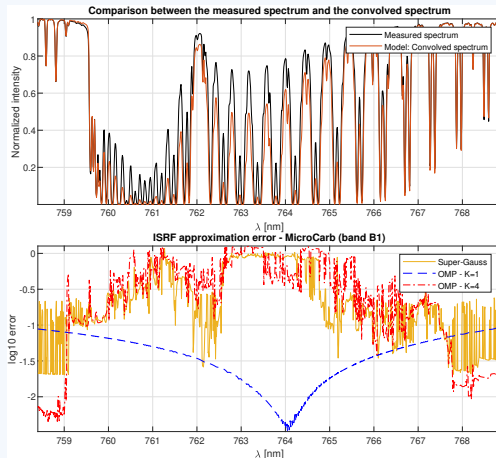
Approximation error MicroCarb (band B1)



Mean approximation error MicroCarb (band B1)

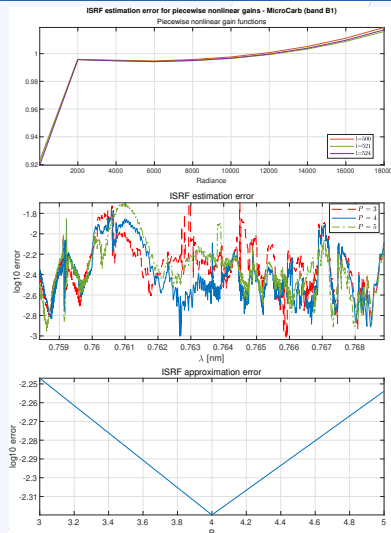


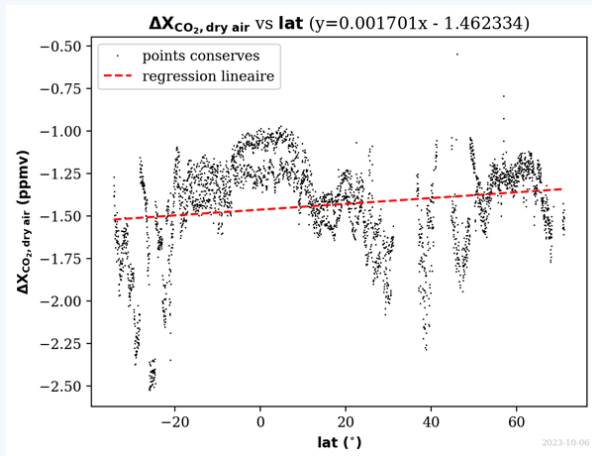
Error in the selection of the reference spectrum



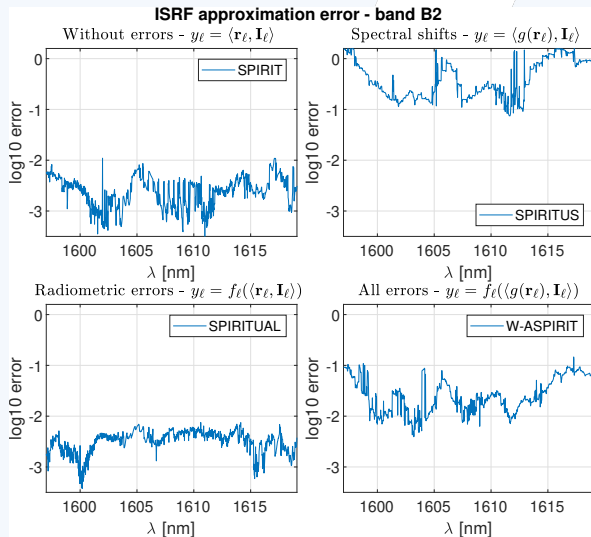
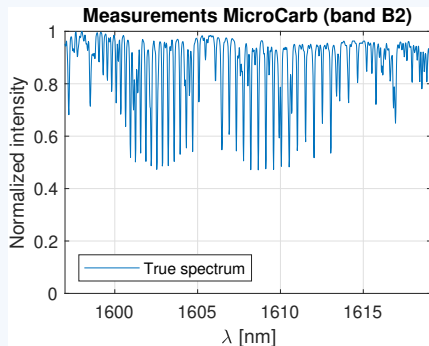
Uncertainties about the nonlinearities

- ▶ $f_I(x) = G_{lin_I}(a_I x + b_I)(a_I x + b_I)$
 with
 - ▶ G_{lin_I} piecewise nonlinear gain
 - ▶ a_I multiplicative gain
 - ▶ b_I additive gain
- ▶ Sparsity $K = 4$

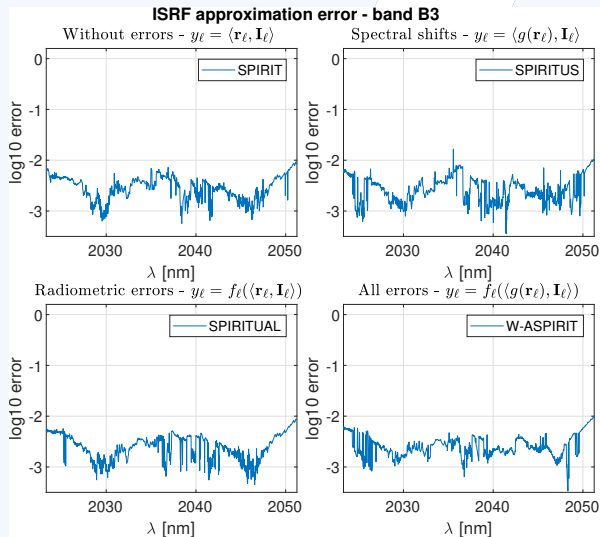
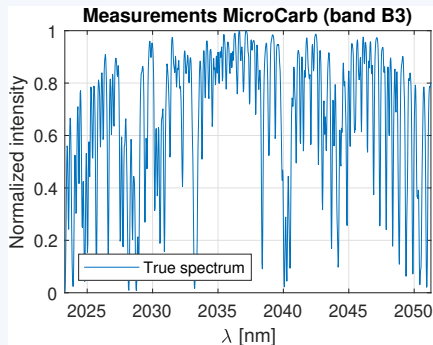




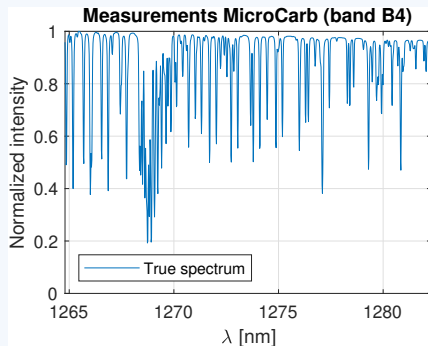
Results for all scenarios and all bands



Results for all scenarios and all bands



Results for all scenarios and all bands

**ISRF approximation error - band B4**